Bayesian TV Denoising of SAR images

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Abstract—Synthetic aperture radar (SAR) imagery suffers from the speckle phenomenon. Speckle gives rise to the presence of multiplicative noise which severely degrades the observed images. It is known that logarithmically transformed speckle can be well approximated by a Gaussian distribution. In this paper we propose an algorithm for despeckling images, within the log-transformed spatial domain, using a TV prior whose model parameter is automatically determined using the Evidence Analysis within the Hierarchical Bayesian Paradigm. The effectiveness of the proposed algorithm, over both synthetically speckled and real SAR images, is studied.

Index Terms—SAR images denoising, despeckling, parameter estimation, Bayesian methods, image restoration.

I. INTRODUCTION

Synthetic aperture radar (SAR) satellite and airborne sensors gather a huge amount of all time, all weather, valuable information. It is well known (see [1]) that SAR, as other coherent imagery, like ultrasound and laser scans, suffer from the speckle phenomenon. Speckle is due to the interferences of waves reflected by many elementary reflectors inside a resolution cell, which, when fully developed, give rise to the presence of multiplicative noise. Fully developed speckle appears when the scanned surface is extremely rough compared to the beam wavelength. In this paper we only consider fully developed speckle. Speckle seriously damages the usability of observed images. A way to mitigate speckle is to take many samples per resolution cell, which is known as multi-look, but at the cost of a loss of resolution.

Many approaches to the problem of denoising speckled images, also known as despeckling, have been proposed. Median and adaptive filters, based on local statistics like the method in [2], have been applied (see [3]). The problem has also been studied within the wavelet and other transformed domains (see [4]–[6]), and several *Total Variation* (TV) applications to despeckling have appeared (see [6]–[9]). Some methods work directly with the observed images [2], [3], [7], [9], while others first apply a log-transform to the images, in order to convert multiplicative into additive noise [4]–[6], [8].

In this paper we propose a simple algorithm for despeckling images, within the log-trasformed spatial domain, using a TV prior. The model parameter is automatically determined using the Evidence Analysis within the Hierarchical Bayesian Paradigm [10].

The rest of this paper is organized as follows. Section II provides the mathematical model for speckled images acquisition process. We provide the description of the hierarchical Bayesian framework modeling the unknowns in Section III. The inference procedure utilized to develop the proposed method is presented in Section IV. We demonstrate the effectiveness of the proposed approach with experimental results in Section V. Finally, section VI concludes the paper.

II. PROBLEM FORMULATION

The degradation model for an *L*-looks observed image g, of size $p = s \times t$, with elements g_i subject to fully developed speckle noise, from the noise free original image f, of the same size, with elements f_i , is the following

$$g_i = f_i \mu_i$$
, for $i = 1, ..., p$. (1)

Therefore μ , of size p, with elements μ_i , in Eq. (1), follows the Gamma distribution

$$p(\mu_i) = \Gamma(\mu_i | L, L), \text{ for } i = 1, ..., p,$$
 (2)

with

$$\Gamma(\omega|a,b) = \frac{b^a}{\Gamma(a)} \omega^{a-1} \exp\left[-b\omega\right],\tag{3}$$

for $\omega > 0$, where a > 0 and b > 0 are the shape and inverse scale parameters, respectively. Its mean and variance are $E[\omega] = a/b$, $Var[\omega] = a/b^2$. So, for $p(\mu_i)$ in Eq. (1) we have $E[\mu_i] = 1$ and $Var[\mu_i] = \frac{1}{L}$.

It has been shown (see [11]) that logarithmically transformed speckle noise approximately follows the Gaussian distribution, that is,

$$\ln \mu_i \sim \mathcal{N}(\mu_i | m, \sigma^2) \tag{4}$$

with mean and variance

$$n = \psi(L) - \ln L, \quad \sigma^2 = \psi^{(1)}(L),$$
 (5)

where $\psi(z) = \frac{d}{dz} \ln \Gamma(z)$ is the digamma function, and $\psi^{(1)}(z) = \frac{d}{dz}\psi(z)$ is the polygamma function of order 1.

Let us define the logarithmically transformed images, of size p,

$$y_i = \ln g_i \,, \tag{6}$$

and

$$x_i = \ln(f_i) + m$$
, for $i = 1, ..., p$, (7)

where m has been defined in Eq. (5). From Eqs. (1) and (4), we obtain the following degradation model, for the 'observed' y from the unknown x,

$$\boldsymbol{y} = \boldsymbol{x} + \boldsymbol{\nu} \,, \tag{8}$$

with $\boldsymbol{y}, \boldsymbol{x}$ and $\boldsymbol{\nu}$, lexicographically arranged vectors of size $p, \boldsymbol{\nu} \sim \mathcal{N}(\boldsymbol{\nu}|\mathbf{0}, \sigma^2 \mathbf{I}), \sigma^2$ defined in Eq. (5), and \mathbf{I} the $p \times p$ identity matrix. Then to obtain an actimate of \boldsymbol{f} from one of $\boldsymbol{\pi}$ one simply used

Then to obtain an estimate of f from one of x one simply uses

$$\hat{f}_i = \exp(\hat{x}_i - m), \text{ for } i = 1, ..., p,$$
(9)

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where \hat{x}_i is an estimate of x_i in Eq. (8) which is obtained with a a denoising method (see [12]).

In this paper the estimate of \hat{x} is obtained by utilizing a TV prior on x. The unknown parameter of this prior distribution is estimated by making use of the Evidence Analysis within the Bayesian framework. Chambolle's algorithm [13] is used to find the maximum '*a posteriori*'estimate of x for a TV prior.

III. BAYESIAN MODELING

Our problem is now stated in terms of a Bayesian inference about the unknown x from the transformed observed image y. We will solve for x within the Hierarchical Bayesian Paradigm (see [10]).

In the hierarchical approach to image restoration we have at least two stages. In the first stage, knowledge about the structural form of the noise and the structural behavior of the restoration is used in forming p(y|x) and $p(x|\alpha)$, respectively.

Given the degradation model defined in Eq. (8), the probability distribution p(y|x) will be

$$p(\boldsymbol{y}|\boldsymbol{x}) = \mathcal{N}(\boldsymbol{y}|\boldsymbol{x}, \sigma^2 \mathbf{I}), \qquad (10)$$

with σ^2 known and defined in Eq. (5).

We will use a TV image prior [14], that is,

$$p(\boldsymbol{x}|\alpha) \propto \alpha^{\frac{p}{2}} \exp\left(-\alpha TV(\boldsymbol{x})\right),$$
 (11)

with the model parameter α , and

$$TV(\boldsymbol{x}) = \sum_{i=1}^{p} \sqrt{\left(\Delta_{i}^{h}(\boldsymbol{x})\right)^{2} + \left(\Delta_{i}^{v}(\boldsymbol{x})\right)^{2}}, \qquad (12)$$

where $\Delta_i^h(\boldsymbol{x})$ and $\Delta_i^v(\boldsymbol{x})$ are the horizontal and vertical first order differences of \boldsymbol{x} .

In the second stage, of the hierarchical approach, we incorporate knowledge about the model parameter α , using the Gamma hyperprior, that is,

$$\mathbf{p}(\alpha) = \Gamma(\alpha | a^o_\alpha, b^o_\alpha) \,, \tag{13}$$

where the Gamma distribution has been defined in Eq. (3).

Finally, the joint probability distribution of our model is

$$p(\boldsymbol{x}, \boldsymbol{y}, \alpha) = p(\boldsymbol{y}|\boldsymbol{x})p(\boldsymbol{x}|\alpha)p(\alpha), \qquad (14)$$

where the distributions $p(\boldsymbol{y}|\boldsymbol{x})$, $p(\boldsymbol{x}|\alpha)$ and $p(\alpha)$ have been defined in Eqs. (10), (11), and (13), respectively.

IV. BAYESIAN INFERENCE

Bayesian inference will be performed in the so-called Evidence Analysis approach [10]. In this approach $\hat{\alpha}$ is first selected as

$$\hat{\alpha} = \underset{\alpha}{\operatorname{argmax}} \int p(\alpha, \boldsymbol{x} | \boldsymbol{y}) \, \mathrm{d}\boldsymbol{x} = \underset{\alpha}{\operatorname{argmax}} \int p(\alpha, \boldsymbol{x}, \boldsymbol{y}) \, \mathrm{d}\boldsymbol{x}$$
$$= \underset{\alpha}{\operatorname{argmax}} \left\{ p(\alpha) \int p(\boldsymbol{y} | \boldsymbol{x}) p(\boldsymbol{x} | \alpha) \, \mathrm{d}\boldsymbol{x} \right\}, \quad (15)$$

and then \hat{x} as

$$\hat{\boldsymbol{x}} = \operatorname{argmax} \left\{ \operatorname{p}(\boldsymbol{y}|\boldsymbol{x}) \operatorname{p}(\boldsymbol{x}|\hat{\alpha}) \right\} .$$
 (16)

Utilizing

$$\mathcal{E}(\boldsymbol{x}, \alpha, \boldsymbol{y}) = \left\{ \frac{1}{2\sigma^2} \parallel \boldsymbol{x} - \boldsymbol{y} \parallel^2 + \alpha \operatorname{TV}(\boldsymbol{x}) \right\}, \quad (17)$$

we have

$$\hat{\alpha} = \operatorname*{argmax}_{\alpha} \left\{ \alpha^{\frac{p}{2}} p(\alpha) \int \exp\left\{ -\mathcal{E}(\boldsymbol{x}, \alpha, \boldsymbol{y}) \right\} d\boldsymbol{x} \right\}$$
(18)

and

$$\hat{\boldsymbol{x}} = \operatorname*{argmin}_{\boldsymbol{x}} \mathcal{E}(\boldsymbol{x}, \hat{\alpha}, \boldsymbol{y}) \,. \tag{19}$$

Unfortunately, we can not directly tackle the minimization in Eq. (18) because of the TV image prior. In earlier work with TV priors (see [15]), this difficulty is overcome by resorting to *majorization*-*minimization* (MM) approaches, which is also the method adopted in this paper.

The main principle of the MM approach is to find a bound of $\mathcal{E}(\boldsymbol{x}, \alpha, \boldsymbol{y})$ which makes the maximization of (18) tractable. Let us first consider the functional

$$M(\boldsymbol{x}, \boldsymbol{u}) = \sum_{i=1}^{p} \frac{\left(\Delta_{i}^{h}(\boldsymbol{x})\right)^{2} + \left(\Delta_{i}^{v}(\boldsymbol{x})\right)^{2} + u_{i}}{2\sqrt{u_{i}}}, \qquad (20)$$

where $\boldsymbol{u} \in (R^+)^p$, is a *p*-dimensional vector that needs to be computed, as it will be shown below. It can be shown that $TV(\boldsymbol{x}) \leq M(\boldsymbol{x}, \boldsymbol{u})$ [15], and this leads to the following upper bound for $\mathcal{E}(\boldsymbol{x}, \alpha, \boldsymbol{y})$

$$\mathcal{E}(\boldsymbol{x}, \alpha, \boldsymbol{y}) \leq \left\{ \frac{1}{2\sigma^2} \| \boldsymbol{x} - \boldsymbol{y} \|^2 + \alpha \operatorname{M}(\boldsymbol{x}, \boldsymbol{u}) \right\} = \mathcal{F}(\boldsymbol{x}, \boldsymbol{u}, \alpha, \boldsymbol{y}),$$
(21)

which results in a lower bound of the integral in Eq. (18), when replacing $\mathcal{E}(\boldsymbol{x}, \alpha, \boldsymbol{y})$ by $\mathcal{F}(\boldsymbol{x}, \boldsymbol{u}, \alpha, \boldsymbol{y})$.

Then the maximization in Eq. (18) is replaced by the maximization of its lower bound, that is, we have

$$\hat{\alpha} \approx \operatorname*{argmax}_{\alpha, \boldsymbol{u}} \left\{ \alpha^{\frac{p}{2}} \mathbf{p}(\alpha) \int \exp\left\{ -\mathcal{F}(\boldsymbol{x}, \boldsymbol{u}, \alpha, \boldsymbol{y}) \right\} \mathrm{d}\boldsymbol{x} \right\} .$$
(22)

Note that alternatively maximizing this bound with respect to the unknowns and the auxiliary variable u results in closer bounds at each iteration. The bound in Eq. (22) is quadratic and therefore it is straightforward to evaluate analytically.

Let us examine the estimation process in Eq. (22) in detail. We can rewrite $\mathcal{F}(\boldsymbol{x}, \boldsymbol{u}, \alpha, \boldsymbol{y})$, in Eq. (21), as

$$\mathcal{F}(\boldsymbol{x}, \boldsymbol{u}, \alpha, \boldsymbol{y}) = \frac{1}{2} \left\{ \frac{1}{\sigma^2} \boldsymbol{y}^t (\boldsymbol{y} - 2\boldsymbol{x}) + \sum_{i=1}^p \alpha \sqrt{u_i} + \boldsymbol{x}^t \mathbf{Q}(\boldsymbol{u}, \alpha) \boldsymbol{x} \right\},$$
(23)

where t denotes the transpose vector, and

$$\mathbf{Q}(\boldsymbol{u},\alpha) = \frac{1}{\sigma^2} \mathbf{I} + \alpha \left(\Delta^{h^t} W(\boldsymbol{u}) \Delta^h + \Delta^{v^t} W(\boldsymbol{u}) \Delta^v \right) .$$
(24)

In the equations above, Δ^d , for d = h, v, represents the $p \times p$ convolution matrix associated with the first order horizontal (or vertical) differences, and W(u) is a $p \times p$ diagonal matrix of the form $W(u) = \text{diag}\left(u^{-\frac{1}{2}}\right)$.

Fixing α and \boldsymbol{u} and expanding $\mathcal{F}(\boldsymbol{x}, \boldsymbol{u}, \alpha, \boldsymbol{y})$ around its minimum on \boldsymbol{x} which we denote by $\bar{\boldsymbol{x}}$, we have

$$\mathcal{F}(\boldsymbol{x}, \boldsymbol{u}, \alpha, \boldsymbol{y}) = \mathcal{F}(\bar{\boldsymbol{x}}, \boldsymbol{u}, \alpha, \boldsymbol{y}) + (\boldsymbol{x} - \bar{\boldsymbol{x}})^t) \mathbf{Q}(\boldsymbol{u}, \alpha) (\boldsymbol{x} - \bar{\boldsymbol{x}}),$$
(25)

and therefore

$$\int \exp\left\{-\mathcal{F}(\boldsymbol{x}, \boldsymbol{u}, \boldsymbol{\alpha}, \boldsymbol{y})\right\} d\boldsymbol{x} = |\mathbf{Q}(\boldsymbol{u}, \boldsymbol{\alpha})|^{-\frac{1}{2}} \times \exp\left\{-\mathcal{F}(\bar{\boldsymbol{x}}, \boldsymbol{u}, \boldsymbol{\alpha}, \boldsymbol{y})\right\}.$$
 (26)

Now, using Eq. (25) in Eq. (22) and differentiating the right hand side of Eq. (22) with respect to α and \boldsymbol{u} we obtain

$$\frac{1}{\hat{\alpha}} = \eta \frac{1}{\alpha^o} + (1 - \eta) \frac{2}{p} \sum_{i=1}^p \sqrt{\hat{u}_i} \,, \tag{27}$$

with $\alpha^o = \frac{a_{\alpha}^o - 1}{b_{\alpha}^o} = \mathbf{Mode}(\mathbf{p}(\alpha))$, where the hyperprior $\mathbf{p}(\alpha)$ has been defined in Eq. (13), and $\eta \in [0, 1)$ is $\eta = \frac{a_{\alpha}^o - 1}{p/2 + a_{\alpha}^o - 1}$. In Eq. (27)

$$\hat{u}_{i} = (\Delta_{i}^{h}(\bar{\boldsymbol{x}}))^{2} + (\Delta_{i}^{v}(\bar{\boldsymbol{x}}))^{2} + \operatorname{tr} \left[\mathbf{Q}(\hat{\boldsymbol{u}}, \alpha)^{-1} \left(\Delta^{h^{t}} \delta(i) \Delta^{h} + \Delta^{vt} \delta(i) \Delta^{v} \right) \right], \quad (28)$$

where $\delta(i)$ is a $p \times p$ matrix with all its values equal to 0, except $\delta(i)_{i,i} = 1$. An approximation for the trace in the above equation is found by using

$$\mathbf{Q}(\boldsymbol{u},\alpha) \approx \bar{\mathbf{Q}}(\boldsymbol{u},\alpha) = \frac{1}{\sigma^2} \mathbf{I} + \alpha z(\boldsymbol{u}) \left(\Delta^{h^t} \Delta^h + \Delta^{vt} \Delta^v \right) ,$$
(29)

with $z(\boldsymbol{u}) = \frac{1}{p} \sum_{i} \frac{1}{\sqrt{u_i}}, \forall \boldsymbol{u} \in (R^+)^p$. With this last approximation in mind, and substituting $\bar{\boldsymbol{x}}$ by $\hat{\boldsymbol{x}}$ defined in Eq.(19), Eq. (28) becomes

$$\hat{u}_i = \left(\Delta_i^h(\hat{\boldsymbol{x}})\right)^2 + \left(\Delta_i^v(\hat{\boldsymbol{x}})\right)^2 + d(\hat{\boldsymbol{u}}, \hat{\alpha}), \qquad (30)$$

with $d(\boldsymbol{u}, \alpha) = \frac{1}{p} \operatorname{tr} \left[\bar{\mathbf{Q}}(\boldsymbol{u}, \alpha)^{-1} \left(\Delta^{h^{t}} \Delta^{h} + \Delta^{vt} \Delta^{v} \right) \right]$. Notice that the substitution of $\bar{\boldsymbol{x}}$ by $\hat{\boldsymbol{x}}$ implies that the expansion in Eq. (25) is carried out around the maximum a posteriori solution defined in Eq. (19), and that the first order term in \hat{x} is close to the one in \bar{x} which is zero. This expansion is expected to produce a better approximation to the solution of Eq. (18) than when the quadratic approximation in Eq. (22) is utilized. Notice also that \hat{x} can be easily calculated using Cambolle's algorithm.

Finally the proposed iterative method is summarized in Algorithm 1.

Algorithm 1 Bayesian TV denoising of speckled images

- Obtain the log transformed observed image $y = \ln g$, set n = 1, $\hat{\alpha}^{n-1} = \alpha^o = \frac{p}{2\text{TV}(y)}$, $u^{n-1} = 0$, and assign a value to η while convergence criterion is not met do
 - 1. Compute $\hat{\boldsymbol{x}}^n = \operatorname{argmin}_{\boldsymbol{x}} \mathcal{E}(\boldsymbol{x}, \hat{\alpha}^{n-1}, \boldsymbol{y})$ by applying Cambolle's method.

2. Compute
$$\hat{u}_i^n = d(\hat{\boldsymbol{u}}^{n-1}, \hat{\alpha}^{n-1}) + (\Delta_i^h(\hat{\boldsymbol{x}}^n))^2 + (\Delta_i^v(\hat{\boldsymbol{x}}^n))^2$$
.
3. Compute $\frac{1}{\hat{\alpha}^n} = \eta \frac{1}{\hat{\alpha}^o} + (1-\eta) \frac{2}{p} \sum_{i=1}^p \sqrt{\hat{u}_i^n}$.
4. Set $n = n+1$.

Obtain the image f by applying to \hat{x} the inverse transform of Eq. (9).

V. EXPERIMENTAL RESULTS

A number of experiments have been carried out in order to evaluate the performance of the proposed algorithm 1 (henceforth referred to as LTV), compare with the method in [2], and with a Wiener restoration of the y image, defined in Eq. (6), to which the inverse transform of Eq. (9) is later applied to obtain f_{Wie} (henceforth referred to as LWie).

The proposed algorithm was ran until the criterion $\|\hat{\alpha}^k - \hat{\alpha}^k\|$ $\hat{\alpha}^{k-1} \| / \max(\hat{\alpha}^k, \hat{\alpha}^{k-1}) < 10^{-3}$ was satisfied, or a number of 10 iterations reached, with $\hat{\alpha}^k$ denoting the $\hat{\alpha}$ estimate at the k iteration step. We have used for η , defined in Eq. (27), the value $\eta = 1 - 0.8/L$, which prevents an experimentally observed trend, of the proposed

TABLE I VALUES OF PSNR AND SSIM FOR DIFFERENT RESTORATIONS OF THE SYNTHETIC IMAGE SET. FOR DIFFERENT NUMBER OF LOOKS L MULTIPLICATIVE NOISE.

	[2]		LWie		LTV		
Image	PSNR	SSIM	PSNR	SSIM	PSNR	SSIM	L
	23.11	0.4922	14.05	0.1777	23.45	0.5477	4
Fig. 1(a)	26.60	0.6694	20.89	0.4494	26.90	0.6909	16
	27.54	0.7169	24.00	0.6028	28.28	0.7425	32
	28.15	0.7460	27.07	0.7404	29.95	0.8014	64
	17.69	0.3336	9.67	0.1172	18.32	0.4192	4
Fig. 1(f)	22.63	0.5486	15.34	0.2794	23.53	0.6437	16
	24.29	0.6465	18.05	0.3818	25.27	0.7259	32
	25.47	0.7251	20.91	0.4941	27.03	0.7954	64
	20.97	0.3616	11.34	0.1019	22.45	0.5203	4
Lena	26.42	0.5785	17.71	0.2549	28.58	0.7638	16
	28.68	0.6798	20.71	0.3542	30.53	0.8207	32
	30.50	0.7638	23.70	0.4649	32.27	0.8610	64
	20.08	0.4186	11.10	0.2432	20.94	0.5173	4
camera	23.48	0.5563	17.40	0.4015	26.01	0.7214	16
1	24.57	0.6380	20.53	0.4790	28.03	0.7912	32
	25.21	0.7096	23.60	0.5569	29.99	0.8429	64

algorithm 1, to obtain too high values for α when the value of the number of looks L grows, producing an extra-smoothing effect.

We first deal with synthetic images. Two aerial images, a $1024 \times$ 1024 image, a detail of which is shown in Fig. 1(a), and a 512×512 one, a detail of which is shown in Fig. 1(f). The 512×512 Lena and the 256×256 cameraman images have also been used in the experiments. Multiplicative noise, corresponding to different number of looks L, has been generated and the resulting degraded images were restored applying the various methods. Table I shows a numerical comparison of the different methods both in term of the peak signalto-noise ratio (PSNR), and the Structural Similarity Index Measure (SSIM) defined in [16], whose maximal value, corresponding to exactly equal images, is +1. In Fig. 1 details from a 360×360 segment of the original, degraded, and the different restorations, for the two aerial images are shown. Both the numerical and the visual comparison of the restored images support the better performance of the proposed algorithm.

Finally a real TerraSAR-X Enhanced Ellipsoid Corrected (EEC) HH polarization channel image, from near Rome, has been considered. A 256×256 image zone of a 25 looks observation is shown in Fig. 2(a), and its corresponding restoration using the proposed method is shown in Fig. 2(b). In this case the real noise free image is not available in order to perform a numerical evaluation of the restoration quality, but the reduction of speckle in the restored image is visually appreciable.

VI. CONCLUSIONS

A novel algorithm for despeckling images, within the logtransformed spatial domain, using a TV prior whose model parameter is automatically determined using Evidence Analysis within the Hierarchical Bayesian Paradigm, has been proposed in this paper. Both the numerical and the visual evaluation of the quality of the restored images support the effectiveness of the proposed algorithm.

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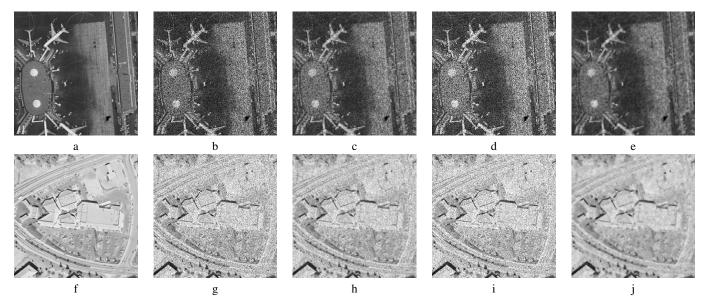
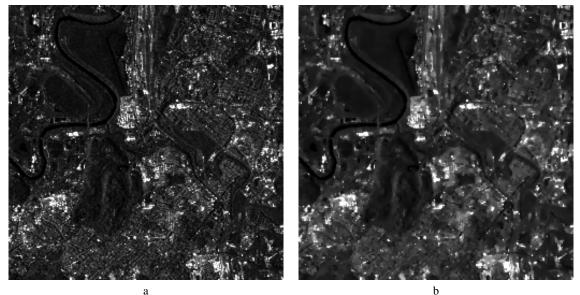


Fig. 1. 360 × 360 details of: (a) & (f) aerial original images, their corresponding degradations for (b) 4 looks and (g) 16 looks, (c) & (h) their restorations using the method in [2], (d) & (i) using LWie and (e) & (j) using the proposed LTV method.



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Fig. 2. (a) 256×256 detail of a EEC HH TerraSAR-X image, and (b) its restoration using the proposed LTV method.

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